

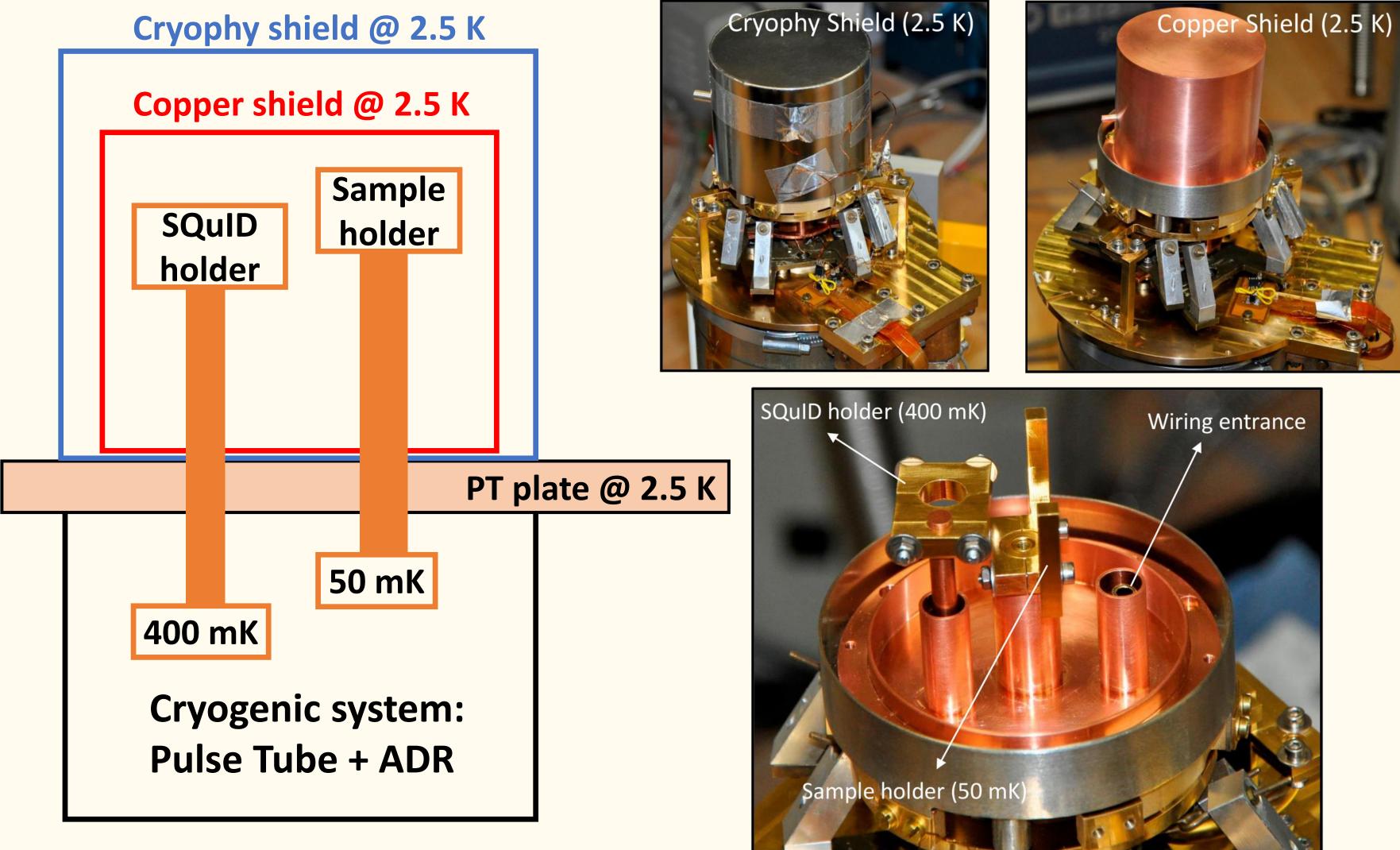
# **The Cryogenic AntiCoincidence detector for ATHENA X-IFU:** improvement of the test setup towards the Demonstration Model

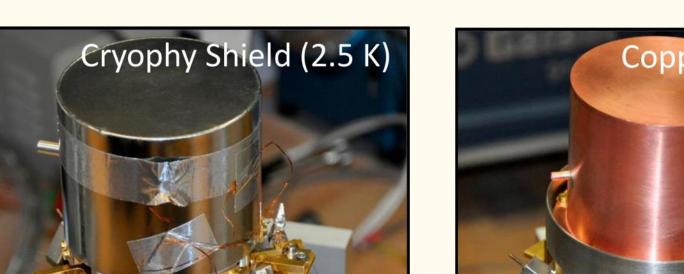


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ATHENA is a large-class ESA mission, to be launched in 2028 towards an L2 orbit. One of the two on-board instruments is the X-ray Integral Field Unit (X-IFU), a cryogenic spectrometer based on a large array of TES microcalorimeters. The X-IFU incorporates a TES-based Cryogenic Anticoincidence detector (CryoAC), placed <1 mm below the main array. The CryoAC development schedule foresees by the end of 2017 the delivery of a Demonstration Model (DM) to the SRON team having in charge the development of the Focal Plane Assembly DM, which will verify some representative detector requirements (see M. Biasotti poster: PE-47). Here we will report the improvement of the cryogenic test setup performed in INAF/IAPS towards the DM test and characterization activities.

## The new cryogenic magnetic shielding system for better quality measurements on CryoAC samples



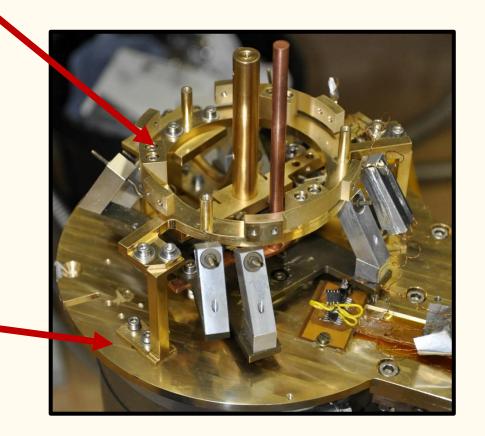


We have upgraded the cryostat at INAF/IAPS inserting a cryogenic magnetic shielding system at the 2.5 K level. The system is constituted by three main parts:

• An external ferromagnetic shield made of Cryophy, a nickeliron soft alloy with high magnetic permeability ( $\mu_r \sim 70000$  at cold), suitable for static and low-frequencies magnetic shielding at cryogenic temperature.

- An internal OFHC copper shield that, if necessary, will be lead plated to operate as a superconducting shield.
- A gold-plated copper support structure designed to compensate the different thermal contractions of the shields.

We have also upgraded the 2.5 K plate of the cryostat, moving from Aluminum to gold-plated Copper, in order to minimize thermal gradients within the plate and do not affect the Pulse Tube performances.



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#### Preliminary performance test with the last CryoAC prototype (AC-S8)

We have tested the cryogenic shielding system integrating in the new setup AC-S8, the last CryoAC single pixel prototype (see M. D'Andrea poster: PE-52).



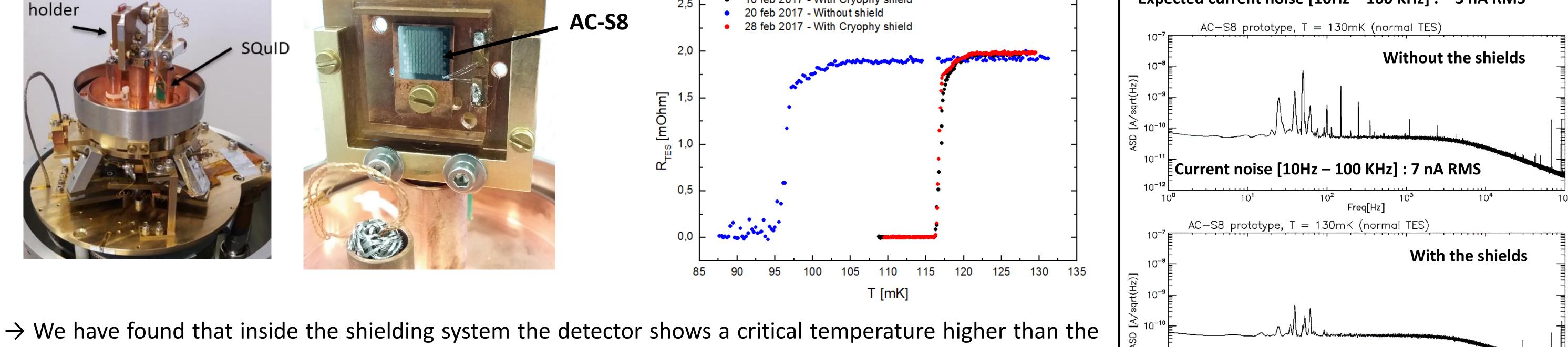


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<ul> <li>10 feb 2017 - With Cryophy shield</li> </ul>																		_

Expected current noise [10Hz – 100 KHz] : ~ 3 nA RMS

Current noise [10Hz – 100 KHz] : 3 nA RMS

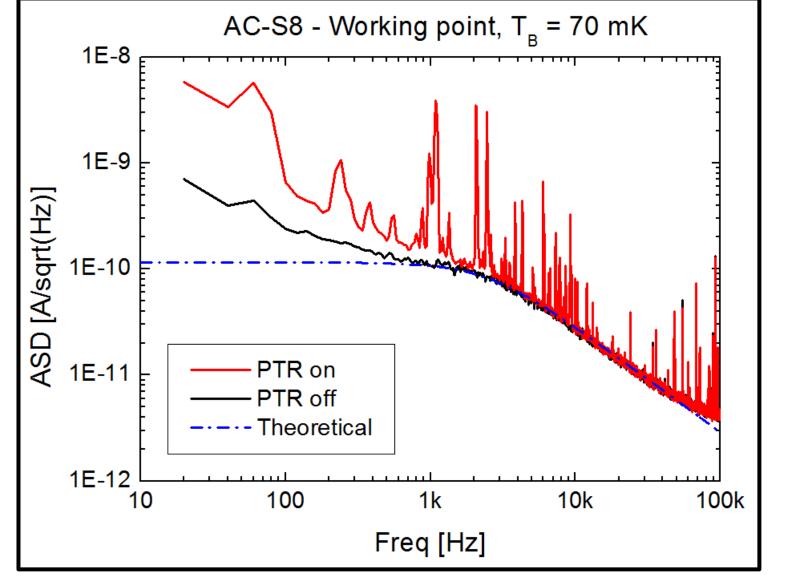
Freq[Hz]



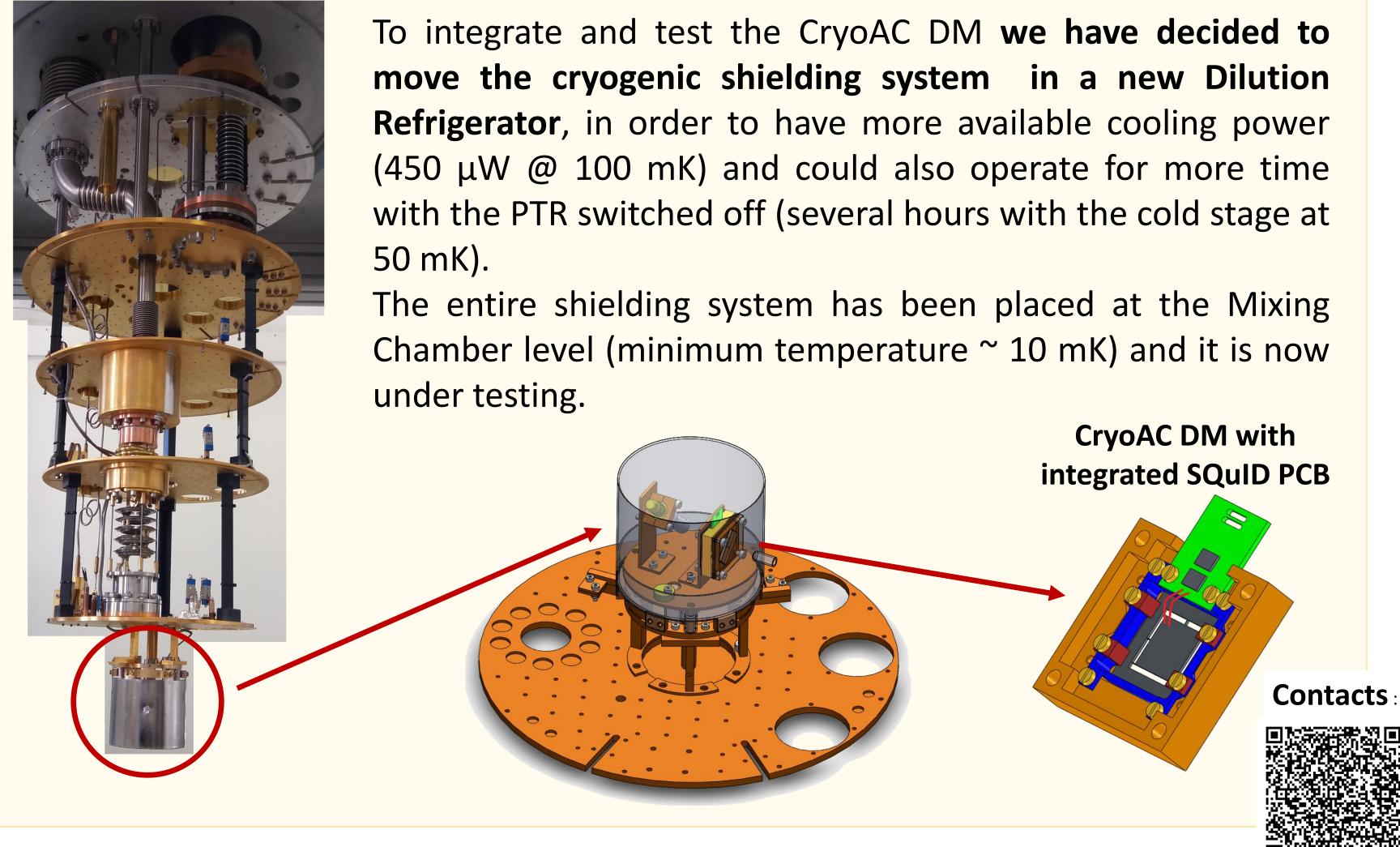
one measured without the shields. This is an evidence of the better magnetic environment.

 $\rightarrow$  The new cryogenic shielding system has also led to a significant reduction in the spectral noise lines.

# The effect of Pulse Tube operation



### The next step: the Demonstration Model integration



 $\rightarrow$  Despite the improvement of the cryogenic setup, we have found that the Pulse Tube Refrigerator (PTR) operation has a dominant impact on the detector noise spectra. This is probably due to the micro-vibration induced on the

cryostat cold stages by the PTR.

 $\rightarrow$  We are able to properly operate the detector **only switching** off the PTR for a small time (typically 15 min with the ADR at 50 mK)