

Optimization of geomagnetic shielding for MKIDs mounted on rotating cryostat

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GroundBIRD is a ground-based CMB experiment which aims to detect the CMB B-modes polarization imprinted by the gravitational waves. We employ Microwave Kinetic Inductance Detectors (MKIDs) as focal plane detectors. We designed magnetic shield for the telescope from experiments and simulations.

Introduction

MKIDs

Microwave Kinetic Inductance Detectors (MKIDs) are superconducting detectors. Below the critical temperature, a photon with energy $h\nu > 2\Delta$ (Δ : gap energy) breaks Cooper pairs, creates quasiparticles (a), and changes surface impedance (b). It results in lower resonance frequency, broader bandwidth (c) and phase shift (d).

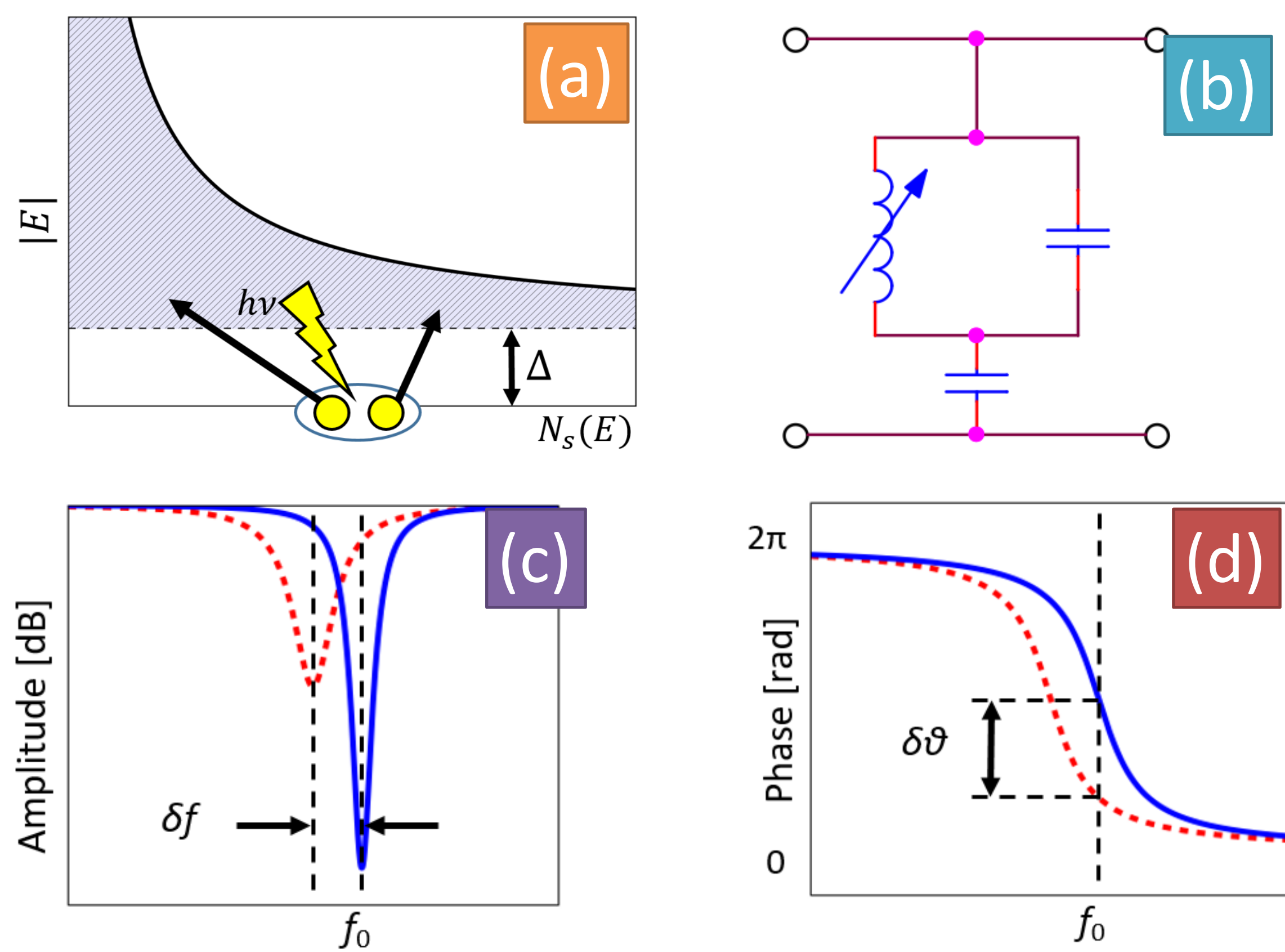


Fig1. Detection principle of MKIDs.

GroundBIRD

One of the key features of the GroundBIRD telescope is the rotational scanning in azimuth direction at 20 rpm to suppress the baseline drift of the detector response caused by 1/f atmospheric fluctuation.

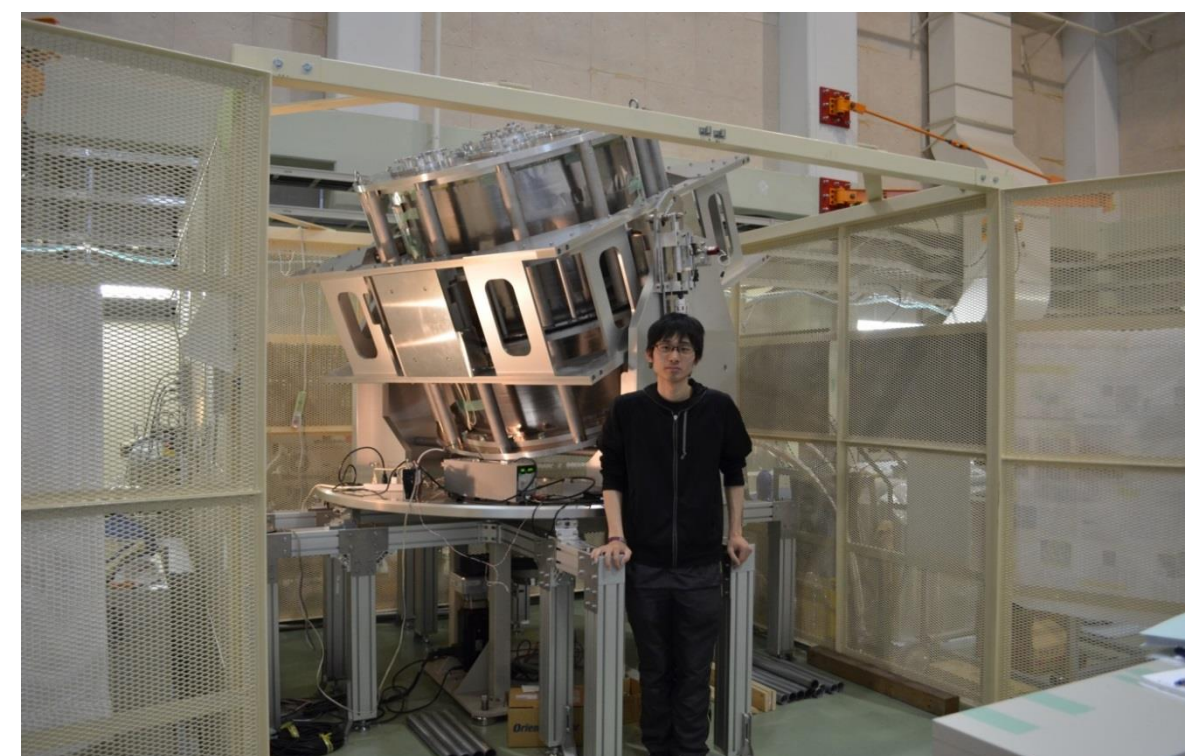


Fig2. GroundBIRD and I.

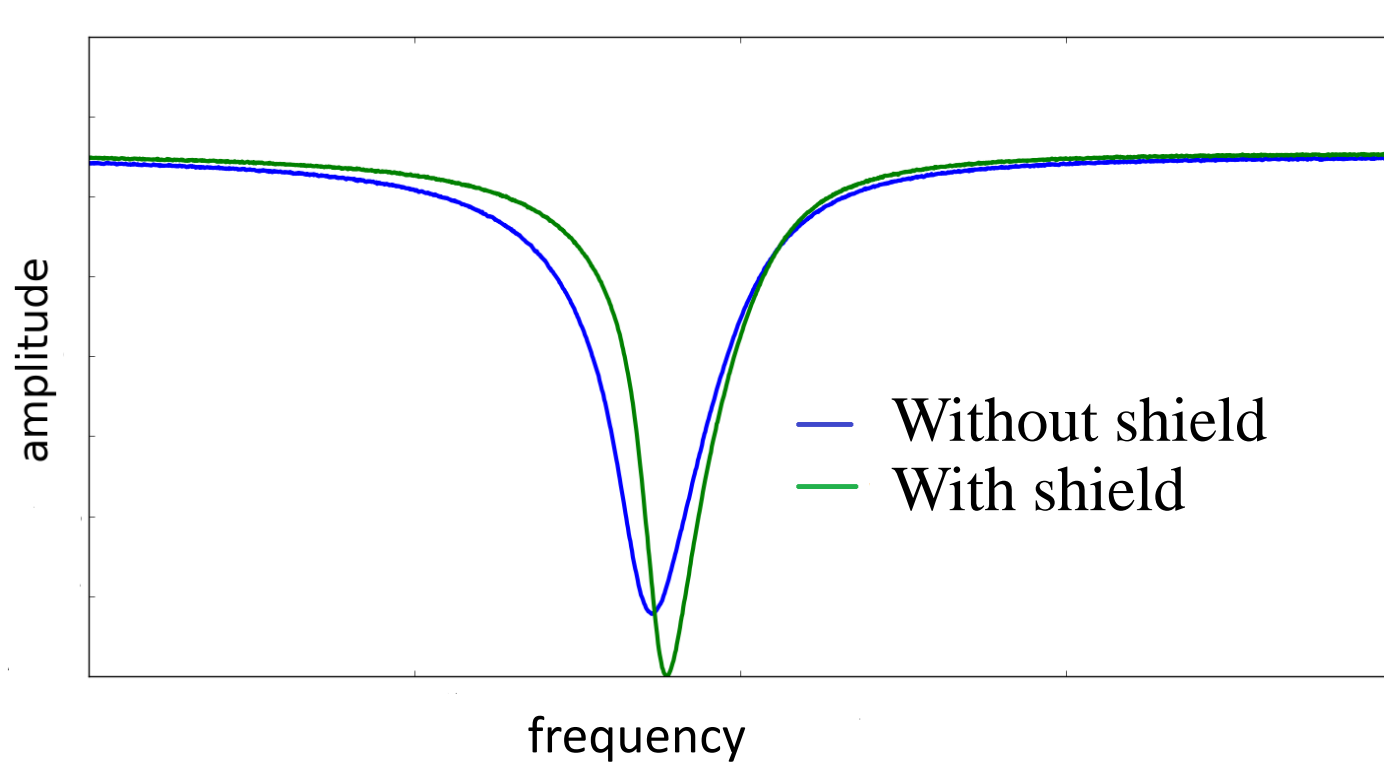


Fig3. Magnetic field effect of MKIDs.

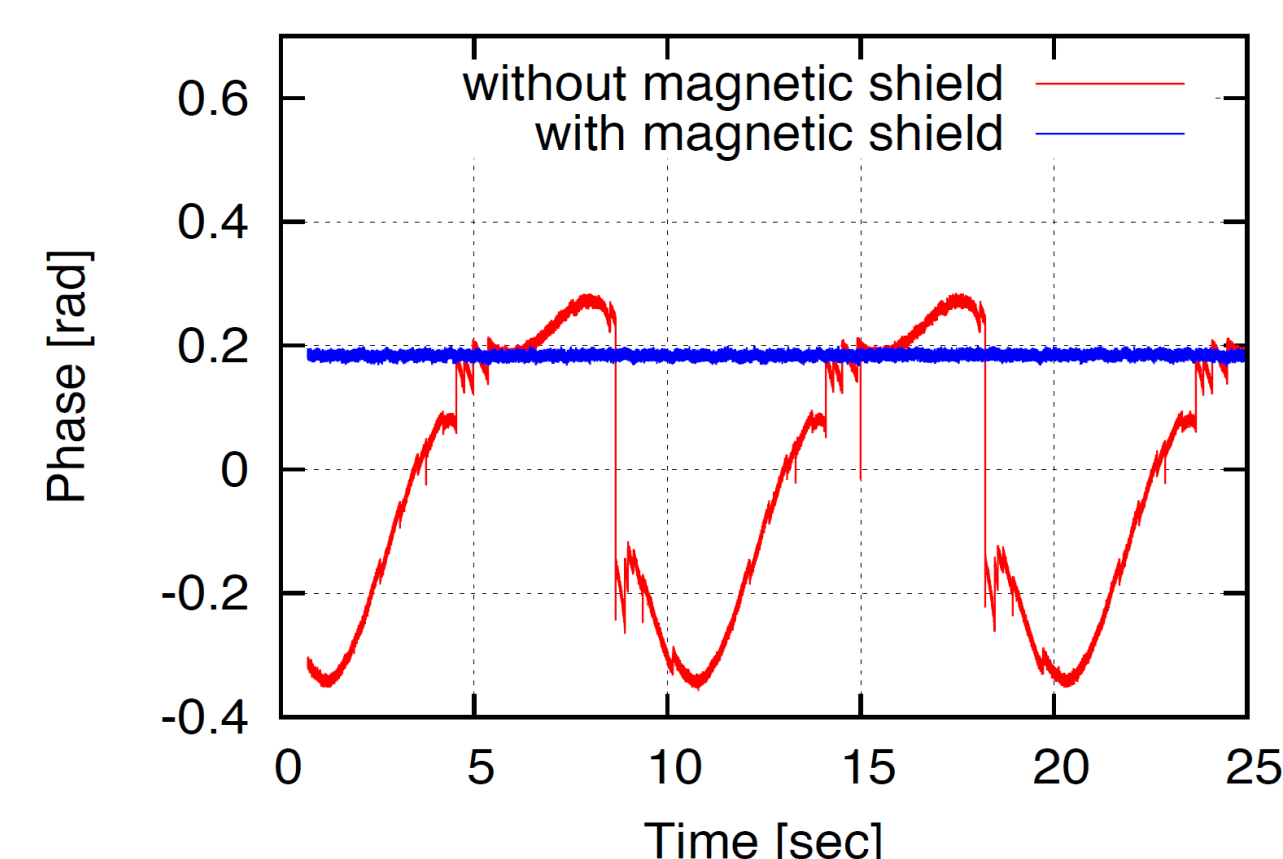
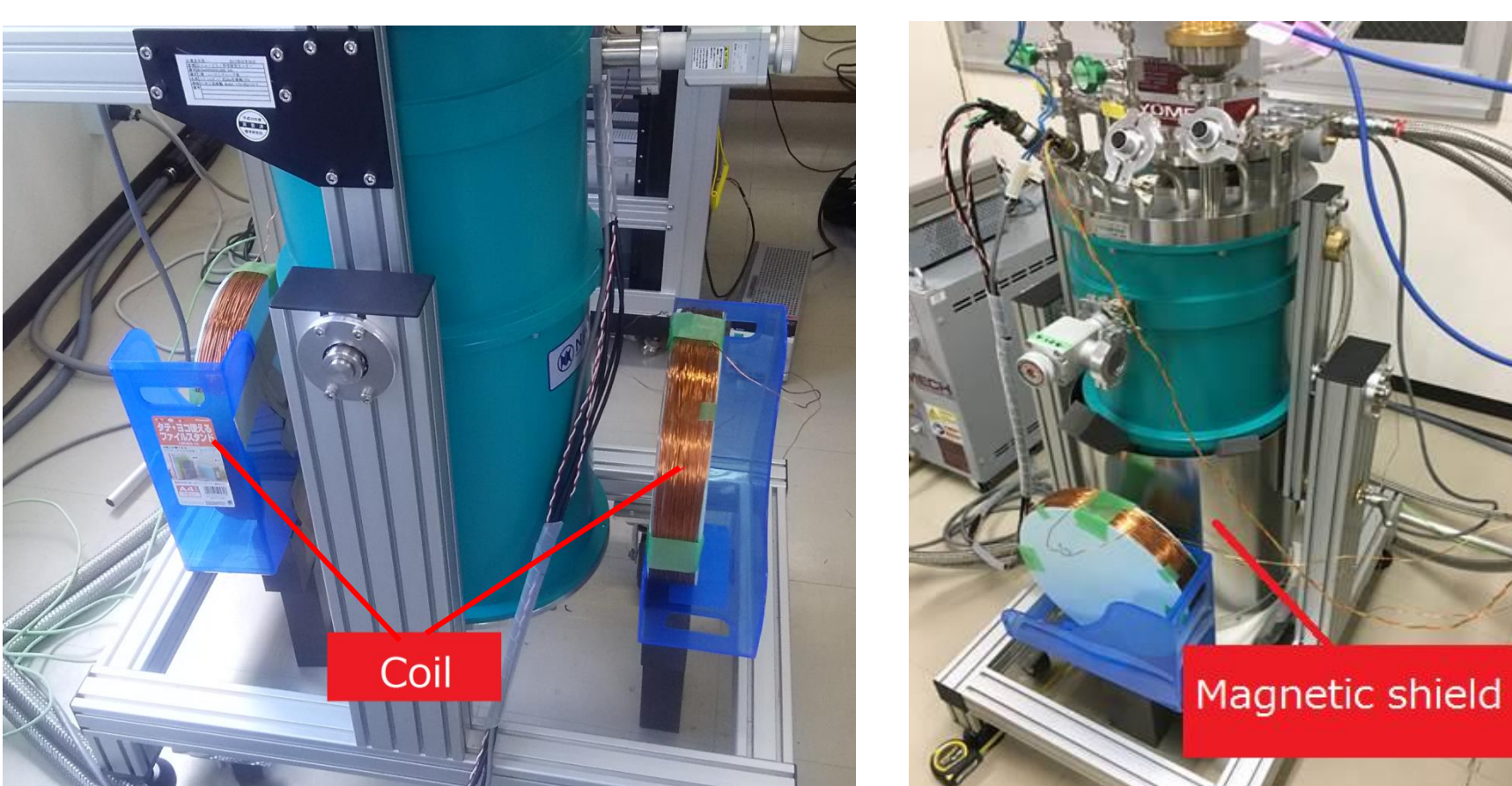


Fig4. Magnetic field effect by rotation.

MKIDs are very sensitive to magnetic fields. For rotating cryostat, change of magnetic field direction causes shift of the center frequency of MKIDs. To solve these problems, we designed a shield to reduce the magnetic field at the focal plane.

Experiment



We measured the magnetic field response of MKIDs using a cryostat in Research Center for Neutrino Science, Tohoku University. All MKIDs are installed in the cryostat. A magnetic shield (40 dB) can be set to the cryostat. We applied magnetic field using a pair of coils and measured the response of MKIDs by Vector Network Analyzer (VNA). We performed measurements (1,2) attaching and detaching the magnetic shield as described in the following table.

Table1. Measurement setup

	Case 1	Case 2
During superconducting transition	NO	YES
During measurement	NO	NO
Δf_r	30kHz	~1kHz

Fig5. Magnetic field Measurement setup.

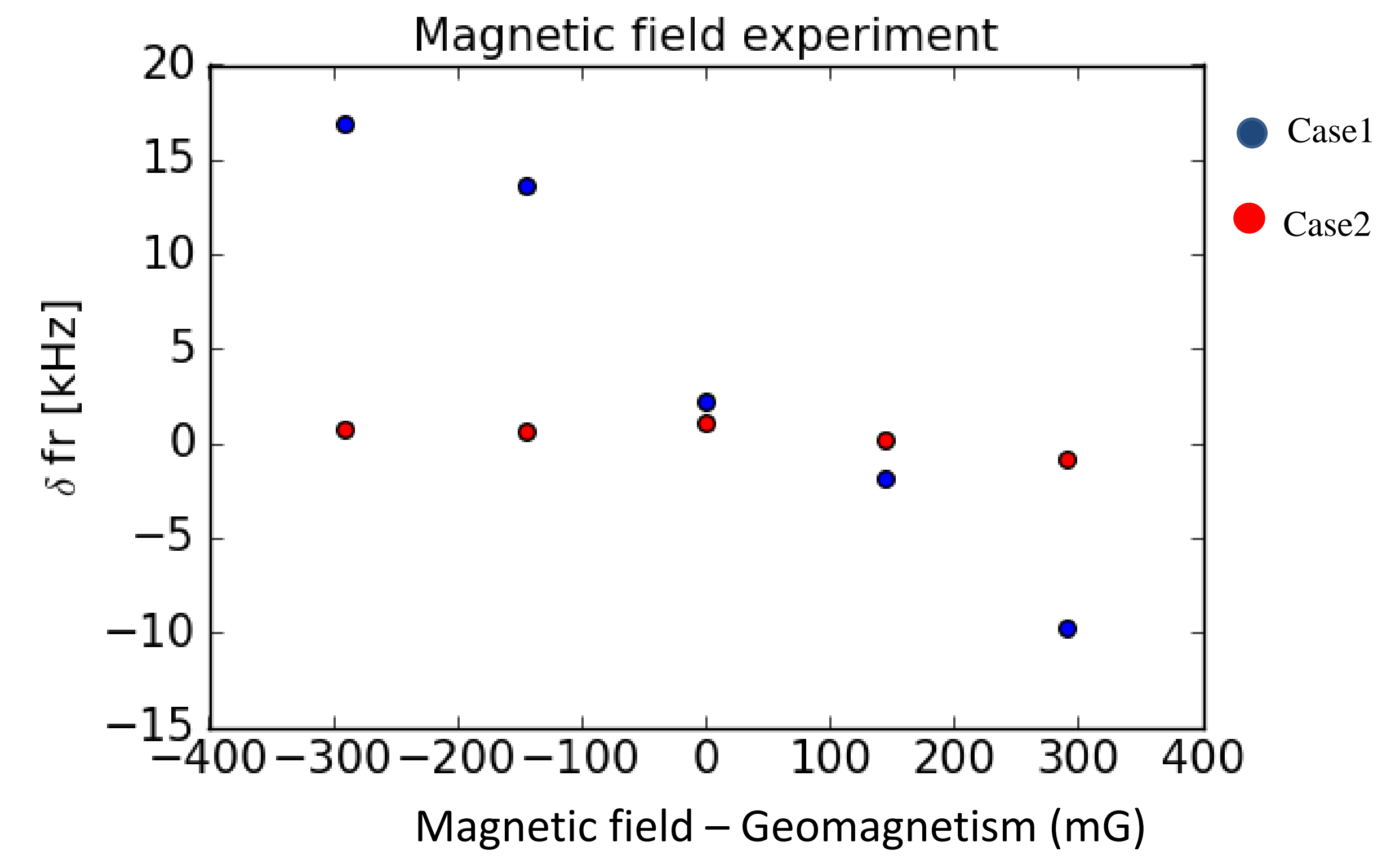


Fig6. The result of magnetic field experiment.

GroundBIRD rotates at 20 rpm, resulting in 1.5-s period variation in the response of MKIDs. Each detector has the frequency range of 125 - 165 GHz. GroundBIRD demands NET ($\sim 300 \mu K\sqrt{s}$) which is converted to power ($P_{GB} \sim 1.2 \times 10^3$ eV/s) in these conditions. Δf_{r1} and Δf_{r2} are also converted to powers (P_1, P_2) using the gap energy (0.18×10^{-3} eV), df_r/dN_{qp} (~ 0.017 Hz) and the time constant τ_{qp} ($\sim 200 \mu s$). We obtain $P_1 = 3 \times 10^6$ eV/s, $P_2 \sim 1 \times 10^5$ eV/s.

If the magnetic shield is set to the cryostat both during the superconducting transition and the experiment (case3), the center frequency shift (Δf_{r3}) is estimated to be less than 0.01 kHz, from which we acquire the equivalent power $P_3 < 1 \times 10^3$ eV/s.

From $P_3 < P_{GB}$, we conclude that about 40 dB magnetic shield is enough for GroundBIRD.

Simulation

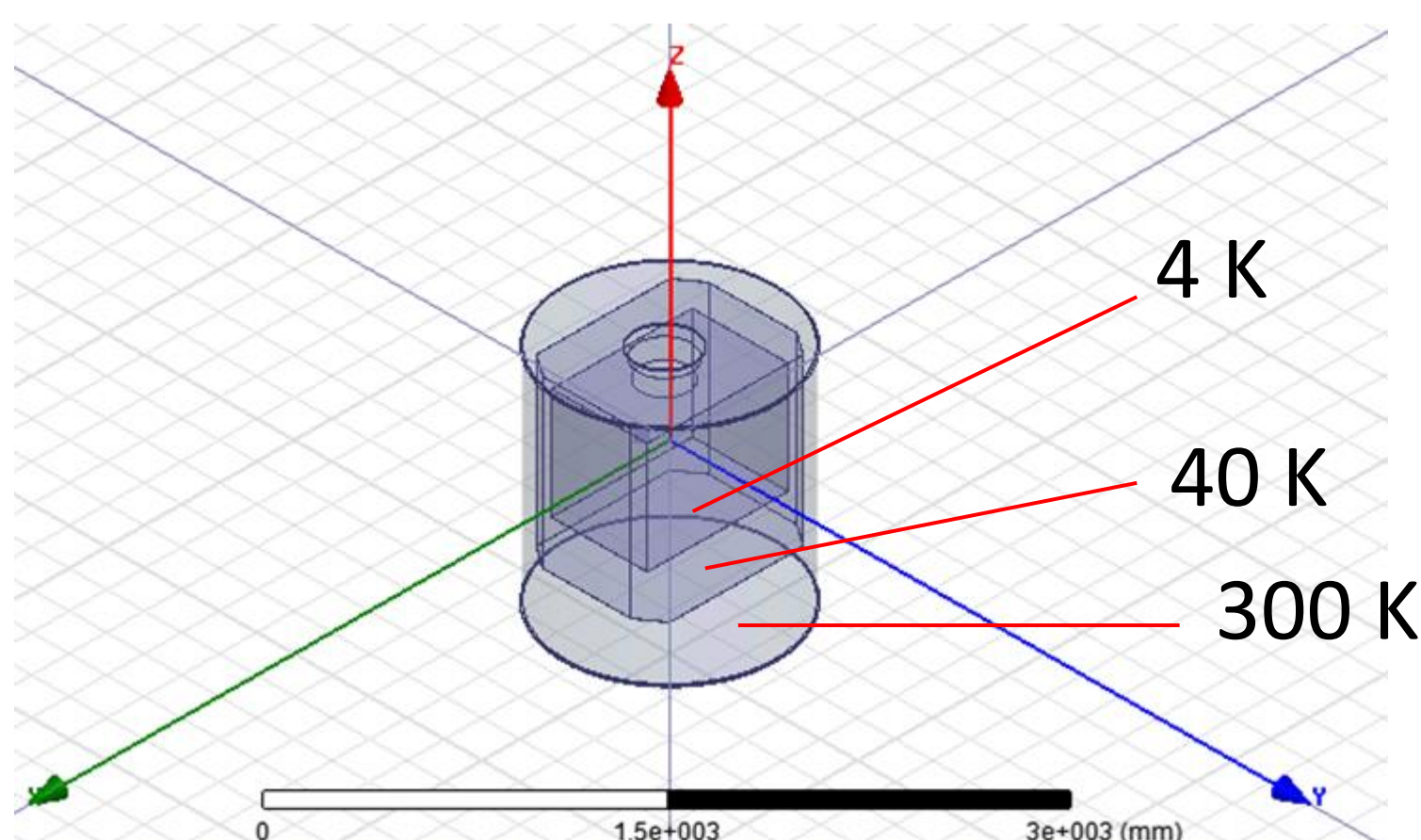


Fig7. Magnetic field simulation.

GroundBIRD has a window to observe sky, and it also allows magnetic flux to enter the cryostat. The window is near detectors on the focal plane and cannot be covered with magnetic shield. The telescope has three cavities (4 K, 40 K, 300 K) on which magnetic shields can be set.

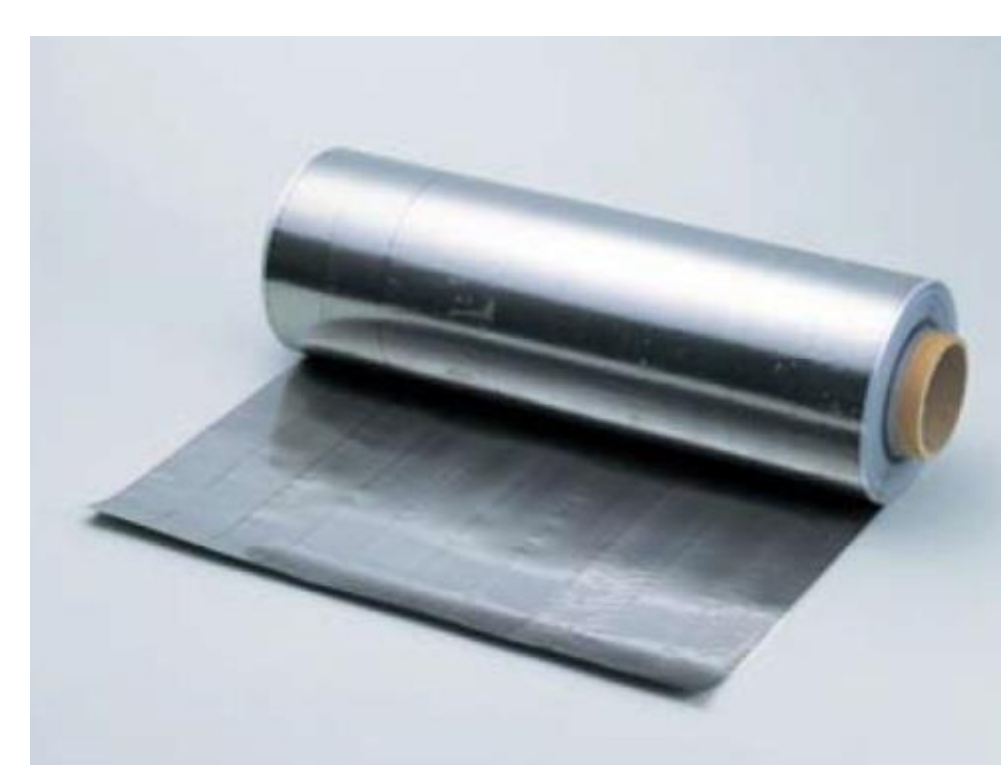


Fig8. Magnetic shield MS-FR.

We used MS-FR made by Hitachi-material as a high permeability magnetic shield. The shield is very thin ($\sim 0.12 \mu m$) and has high permeability ($\mu \sim 70000$).

In order to design the magnetic shield, we evaluated the shielding efficiency by ANSYS Maxwell, which employs finite element method to simulate electromagnetic field.

Shield effect decreases about 5 dB when 5 cm gap is formed in each cavity. From measurement and simulation we conclude that the sensitivity of MKIDs is not affected by geomagnetism if 9 sheets of MS-FR are set on each cavity of the cryostat.

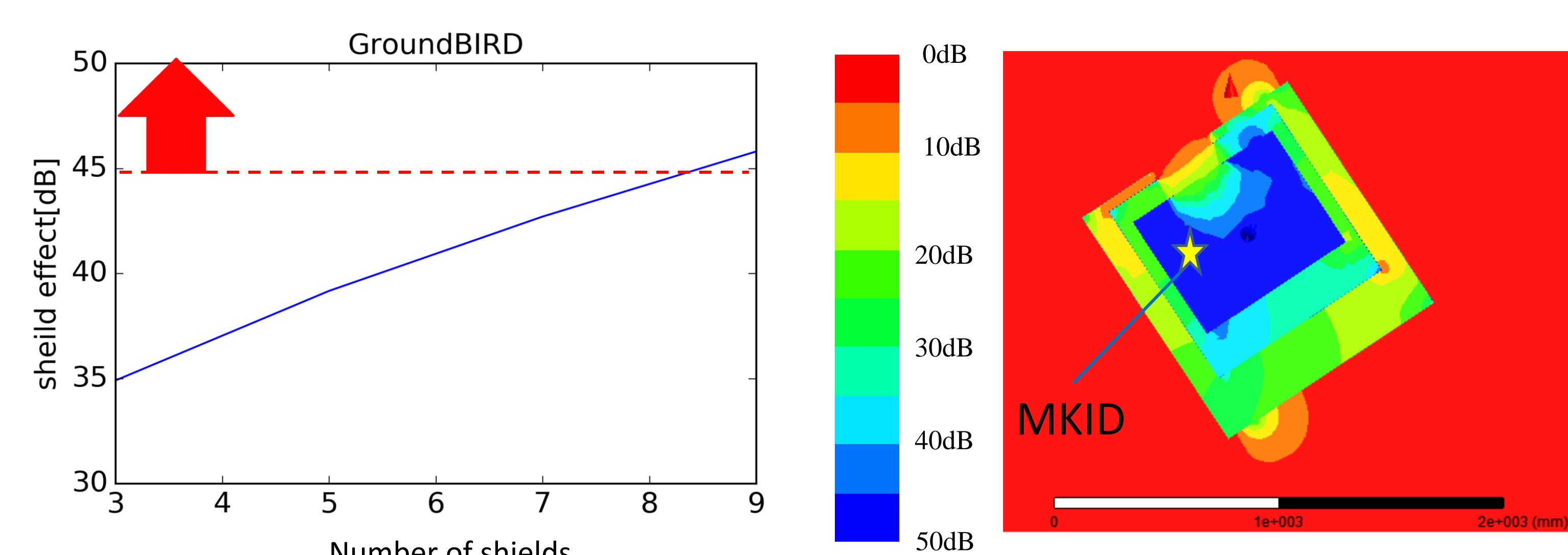


Fig9. Number of magnetic shield vs magnetic shield effect by Maxwell.

Fig10. Maxwell simulation. Over 45 dB suppression is achieved near the focal plane.