

Second-generation design of Micro-Spec: a medium-resolution, submillimeter-wavelength spectrometer-on-a-chip

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would improve capability observatories.

High-altitude provide both the first testbed demonstrate the µ-Spec technology in a space-like economically viable venue for multiple observation campaigns.





Figure 1. Layout of the μ -Spec module. The light is coupled into the instrument via a broadband antenna and is transmitted through a low-loss superconducting transmission line to a divider and a phase delay network. In the multimode region, the feed horns radiate a converging circular wave, which concentrates the power along the focal surface, with different wavelengths at different locations. To disentangle the various orders, the receiver antennas are connected to a bank of order-sorting filters terminated in microwave kinetic inductance detectors (MKIDs) for readout.



The design methodology results in an optimized geometry arrangement for the transmitting and receiving antennas (red arc in Fig. 4) through the minimization of the root-mean-square (RMS) phase error on the focal plane (Fig. 5) [12].

2)

1) INPUT PARAMETER SELECTION

Table 1: Spectrometer design parameters for the configuration selected to fit 4 spectrometers in a 10-cmdiameter silicon wafer. This choice has enabled several designs to be studied as well as the development of the required fabrication process.



Figure 7 (Adapted from [13]). (1) Clean a siliconon-insulator (SOI) wafer. (2) Sputter-deposit the niobium (Nb) ground plane (GP) with argon. (3) Spin-coat bisbenzocyclobutene (BCB) on the Nbcoated surface of the SOI wafer. (4) Manually flip the wafer stack upside down to start processing the SOI wafer backside. (5) Etch the silicon handle wafer by mechanical lapping, followed by deep reactive ion etching using the Bosch process. (6) Deposit molybdenum nitride (Mo₂N), pattern the resonators and sputter-deposit the Nb transmission lines.

References



• In Z-Spec, propagation occurs in parallel-plate waveguides [2].



- [1] H. A. Rowland, "On concave gratings for optical purposes," Philos. Mag. 16, 1883.
- [2] L. Earle et al., "Z-Spec: a broadband, directdetection, millimeter-wave spectrometer instrument status and first results," Proc. SPIE 6275, 2006.
- [3] W. Rotman and R. F. Turner, "Wide-Angle Microwave Lens for Line Source Applications," IEEE T. Antenn. Propag. 11(6), 1963.
- [4] A. Kovács et al., "SuperSpec: design concept and circuit simulations," Proc. SPIE 8452, 2012.
- [5] A. Endo et al., "Design of an Integrated Filterbank for DESHIMA: On-Chip Submillimeter Imaging Spectrograph Based on Superconducting Resonators," J. Low Temp. Phys. 167(3-4), 2012.
- [6] C. N. Thomas et al., "Progress on the CAmbridge Emission Line Surveyor (CAMELS)," 26th Int. Symp. Space THz Technol., 2015.
- [7] S. Bryan et al., "WSPEC: A Waveguide Filter-Bank Focal Plane Array Spectrometer for Millimeter Wave Astronomy and Cosmology," J Low Temp. Phys. 184(1-2), 2016.
- [8] G. Cataldo et al., "Micro-Spec: an ultracompact high-sensitivity spectrometer for far-infrared and submillimeter astronomy," Appl. Opt. 53(6), 2014.
- [9] A. Patel et al., "Fabrication of MKIDs for the Micro-Spec Spectrometer," IEEE T. Appl. Supercon. 23(3), 2013.
- [10] O. Noroozian et al., "µ-Spec: An Efficient Compact Integrated Spectrometer for Submillimeter Astrophysics," 26th Int. Symp. Space THz Technol., 2015. [11] G. Cataldo, "Development of ultracompact, high-sensitivity, space-based instrumentation for far-infrared and submillimeter astronomy," PhD Thesis, MIT, 2015.

- Bootlace lenses are a 1-dimensional analog of Z-Spec [3], which µ-Spec builds on for submillimeter wave applications.
- Narrow-band filter-bank spectrometers do not rely on optical interference as in grating or Fabry-Perot spectrometers. Some examples are: SuperSpec [4, figure below], the Delft SRON High-redshift Mapper (DESHIMA) [5], the CAMbridge Emission Line Surveyor (CAMELS) [6], and similar alternatives made in rectangular waveguides (e.g., W-Spec [7]).



Antenna feed horn design



Figure 6 (Adapted from [8]). (Left) Computed feed horn angular response at 430 GHz. The response of an emitting feed is evaluated in the far field $(r \sim 6\lambda = 438 \ \mu m)$ and normalized to the magnitude of the E field at 0°. The feed array geometry is provided in the figure insert. The feed's phase center is indicated by a black filled circle.

Work is in progress to explore a different design, which employs magnetically coupled antennas. FOR FURTHER DETAILS, PLEASE SEE: B. Bulcha et al., "Electromagnetic Design of a Magnetically-Coupled Spatial Power Combiner," LTD17 #2312245.

- [12] G. Cataldo et al., "Micro-Spec: an integrated direct-detection spectrometer for far-infrared space telescopes," Proc. SPIE 9143, 2014.
- [13] G. Cataldo et al., "Fabrication and Characterization of Superconducting Resonators," J. Vis. Exp. 111, e53868, 2016.

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