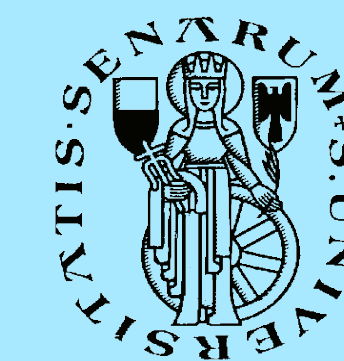


Critical temperature tuning of Titanium thin films for CMB detectors on the SWIPE/LSPE experiment

D. Vaccaro^{a,d}, B. Siri^{b,e}, A.M. Baldini^a, M. Biasotti^e, F. Cei^{a,c}, V. Ceriale^{b,e}, M. de Gerone^b, L. Galli^a, G. Gallucci^b, F. Gatti^{b,e}, M. Grassi^a, D. Grosso^{b,e}, D. Nicolò^{a,c}, M. Piendibene^a, G. Signorelli^a and F. Spinella^a.



^aINFN Sezione di Pisa, Largo B. Pontecorvo 3, 56127 Pisa, Italy
^bINFN Sezione di Genova, Via Dodecaneso 33, 16146 Genova, Italy
^cDipartimento di Fisica, Università di Pisa, Largo B. Pontecorvo 3, 56127 Pisa, Italy
^dDipartimento di Fisica, Università di Siena, Strada Laterina 8, 53100 Siena, Italy
^eDipartimento di Fisica, Università di Genova, Via Dodecaneso 33, 16146 Genova, Italy



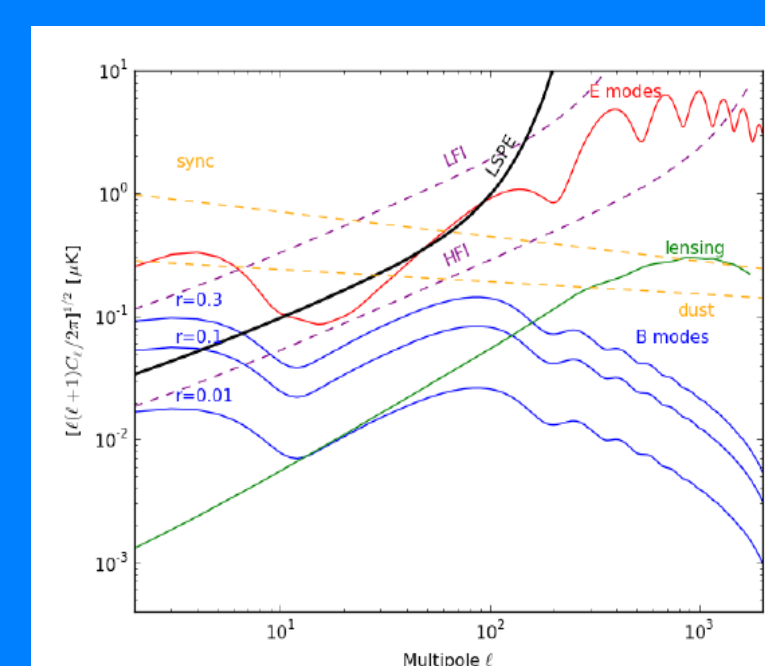
ABSTRACT

Cosmic microwave background (CMB) B-mode polarization detection is a major challenge in modern cosmology, which future experiments are going to undertake either from the ground, on balloons or on satellites. Among these, the SWIPE/LSPE balloon-borne experiment aims at searching for B-modes exploiting the re-ionization peak at large angular scales. Detectors in SWIPE are Transition Edge Sensor (TES) spider-web bolometers, requiring $T_c > 500$ mK. We found evidence that temperature control during deposition and post annealing of Titanium thin films allows the tuning of critical temperature. Titanium is a Type I superconductor with $T_c \approx 390$ mK. In this paper we present a systematic study done on thermal treated Titanium films, showing that higher T_c can be achieved, in a range suitable for SWIPE.

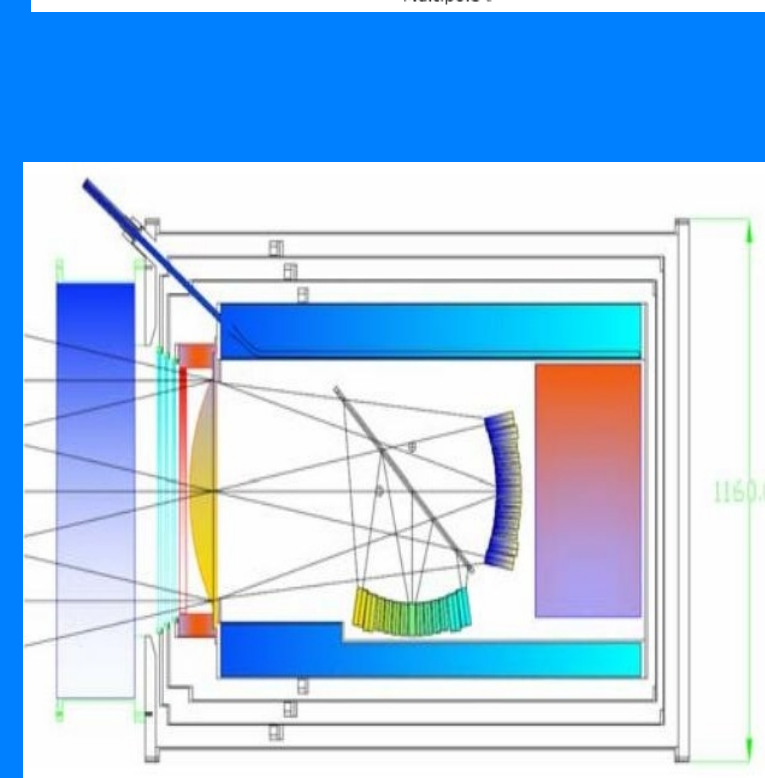
The SWIPE/LSPE experiment

The Large Scale Polarization Explorer¹ (LSPE):

- is a mixed ground/balloon-borne mission aimed at measuring the CMB **B-mode** polarization at large angular scales.
- will improve the limit on the tensor-to-scalar ratio down to $r = 0.03$ at 99.7% CL.
- will be launched from Longyearbyen (Svalbard Islands) in December 2018, for a 15 days-long circumpolar flight during the polar night (40 km of altitude, ≈ -80 °C).
- is composed of two instruments: STRIP (installed on ground at Tenerife) and **SWIPE** (on board).



Upper left: power spectrum of CMB polarization (E-modes and B-modes), foregrounds (synchrotron and interstellar dust) and spurious B-modes due to gravitational lensing. In black is reported the sensitivity limit of LSPE.

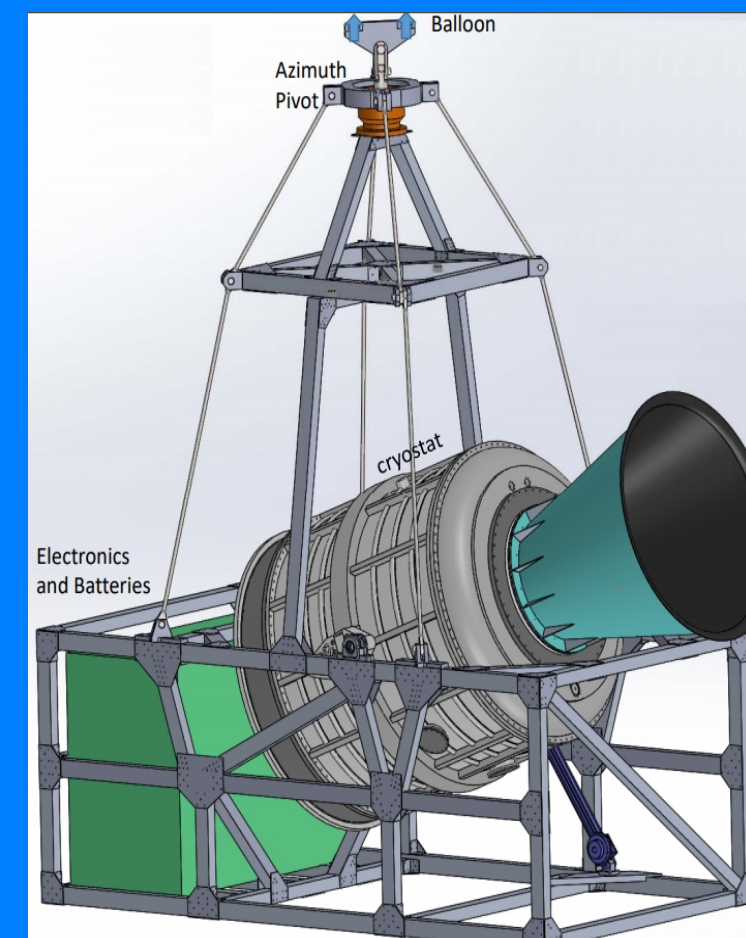


Lower left: sketch of the cryostat of SWIPE/LSPE, hosting the rotating HWP, the beam splitter and the two focal planes cooled at 300 mK.

Lower middle: table with the main (theoretical) parameters for the TESs to be used on SWIPE.

Lower right: one of the multimode horns to be mounted on the focal plane. The dismounted backshot shows the housing for the spiderweb TES.

Frequency (GHz)	140	230	340
χ_{moder}	12	30	34
NEP _{opt} (W/√Hz)	6.4×10^{-17}	5.5×10^{-17}	11.5×10^{-17}
NEP _{th} (W/√Hz)	4.0×10^{-17}	3.2×10^{-17}	5.5×10^{-17}
P_{opt} (pW)	11	5	20
P_{th} (pW)	28	15	50
G (pW/K)	220	120	400



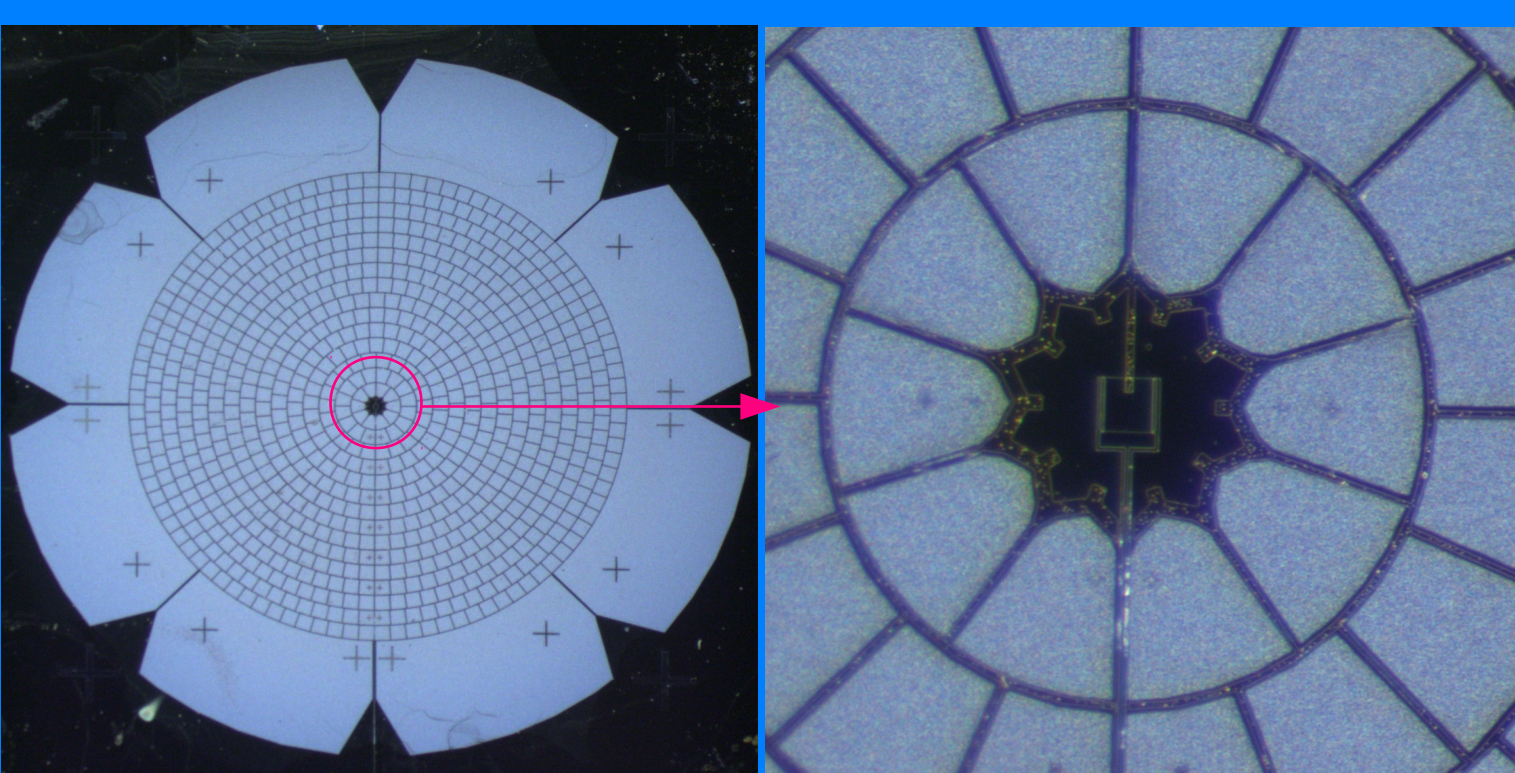
The Short Wavelength Instrument for the Polarization Explorer² (SWIPE):

- is composed of two focal planes with 326 spiderweb TES bolometers cooled at 300 mK.
- exploits multi-mode horns to convey radiation.
- detectors are Titanium Transition Edge Sensors with the requirement of $T_c > 500$ mK, due to the increased optical power collected by the multi-mode horns (see Table³), since at equilibrium:

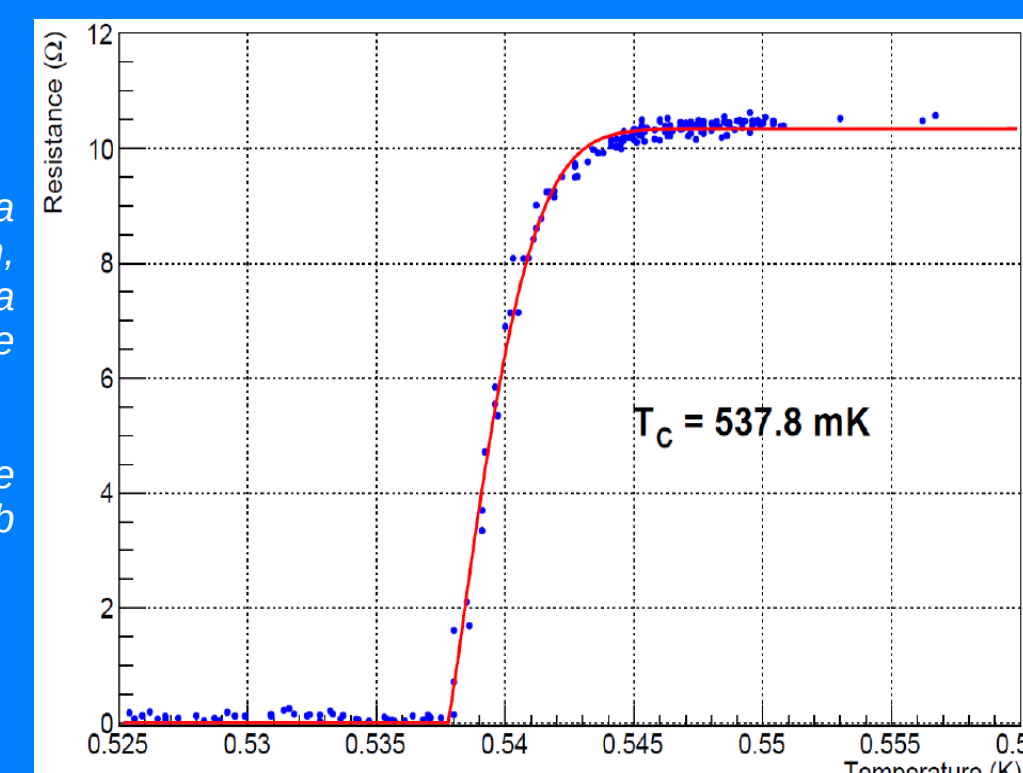
$$T_{\text{eq}} = T_{\text{bath}} + (P_{\text{opt}} + P_{\text{Joule}})/G$$
- readout is achieved by means of a 16-channel FDM, putting a superconducting LC filter in series with each TES. FDM chains are hosted on custom PCBs on the focal planes.
- hot electronics provides comb generation, demodulation and data reduction.

Transition Edge Sensors

- TES bolometers for SWIPE are fabricated at INFN Genova facilities⁴:
- the spiderweb structure of the SiN absorber reduces the energy release from cosmic rays.
 - fabrication requires several steps (resist coating, optical lithography, e-gun and sputtering deposition, ICP-RIE etching, chip release and cleaning).
 - the Titanium film is deposited in a temperature-monitored environment to maintain its T_c above 500 mK, as required.

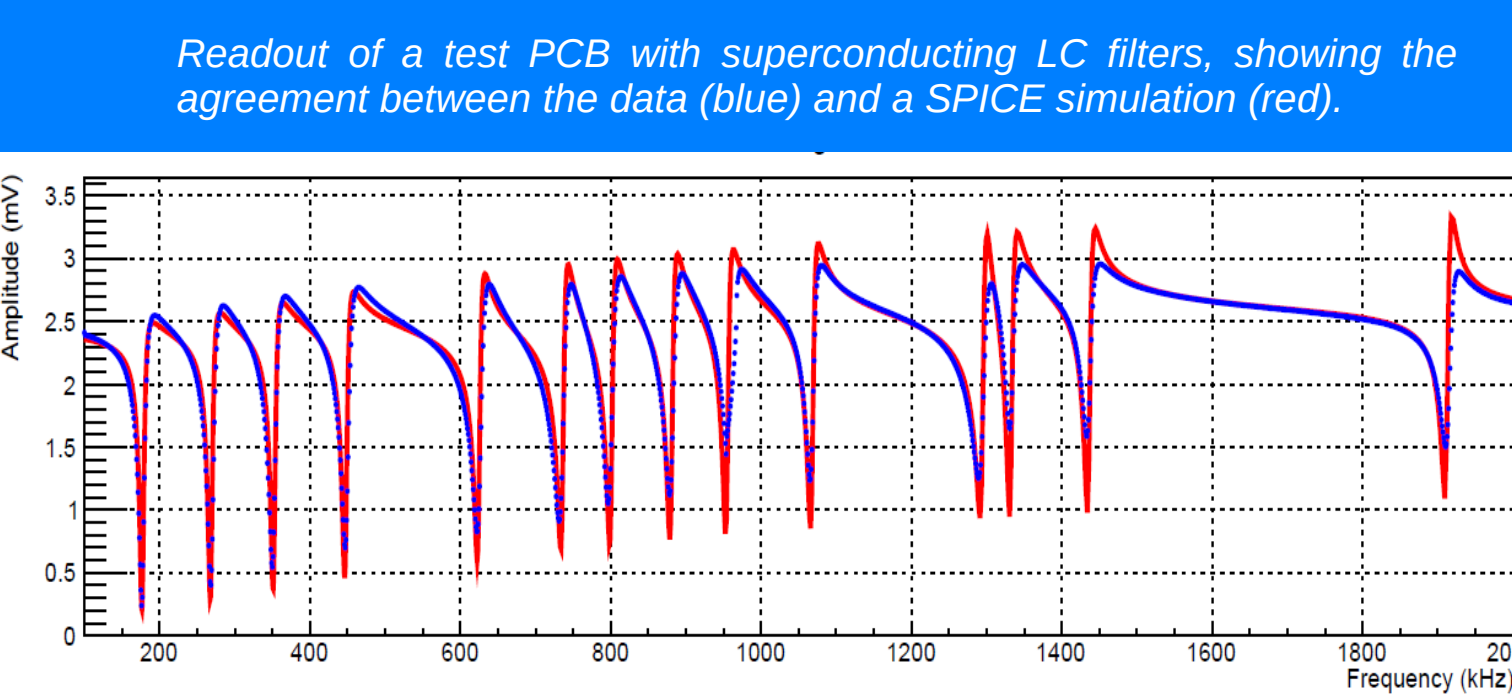


Left: transition of a Titanium film, deposited with a substrate temperature lower than 80 °C. Right: microscope view of a spiderweb TES chip.

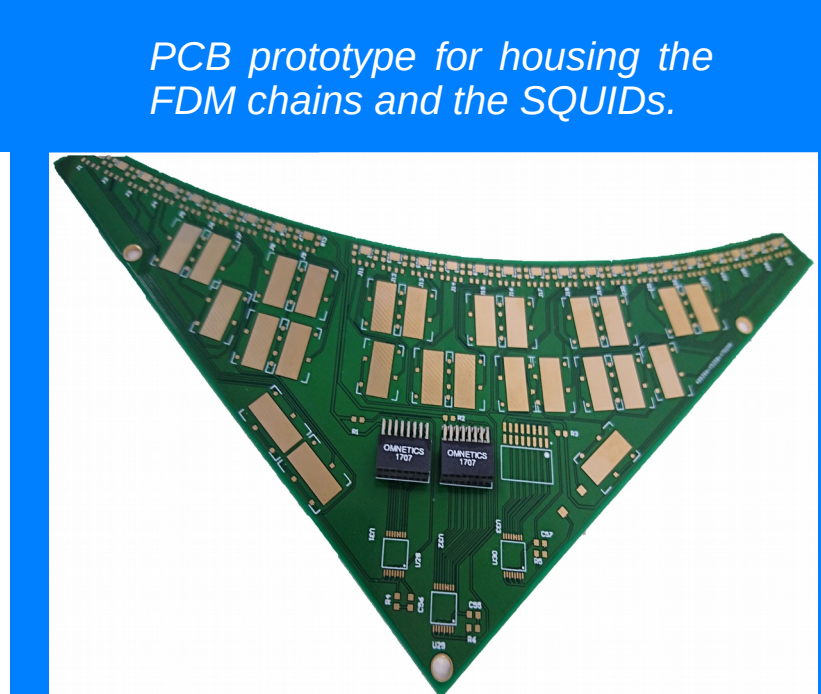


TES detectors readout

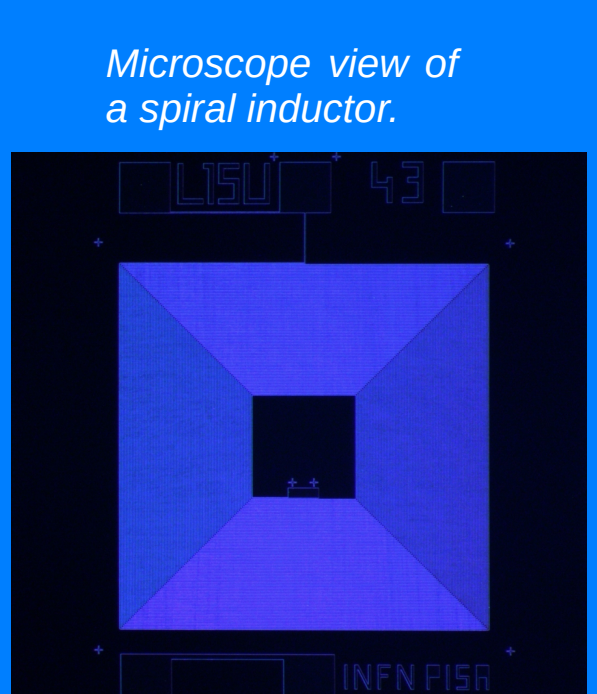
- Readout electronics based on frequency domain multiplexing is being designed, realized and tested at INFN Pisa⁵:
- superconducting LC filters are fabricated at NEST-CNRnano⁶ facilities in Pisa, using optical lithography and Niobium sputtering deposition.
 - testing of the FDM chain is performed at INFN Pisa facilities, including the readout with SQUIDs and hot electronics.
 - first results prove that we can pack 16 resonances in the range 200 kHz ÷ 2 MHz and that inter-channel crosstalk is below 0.3% level⁷.



Readout of a test PCB with superconducting LC filters, showing the agreement between the data (blue) and a SPICE simulation (red).



PCB prototype for housing the FDM chains and the SQUIDs.

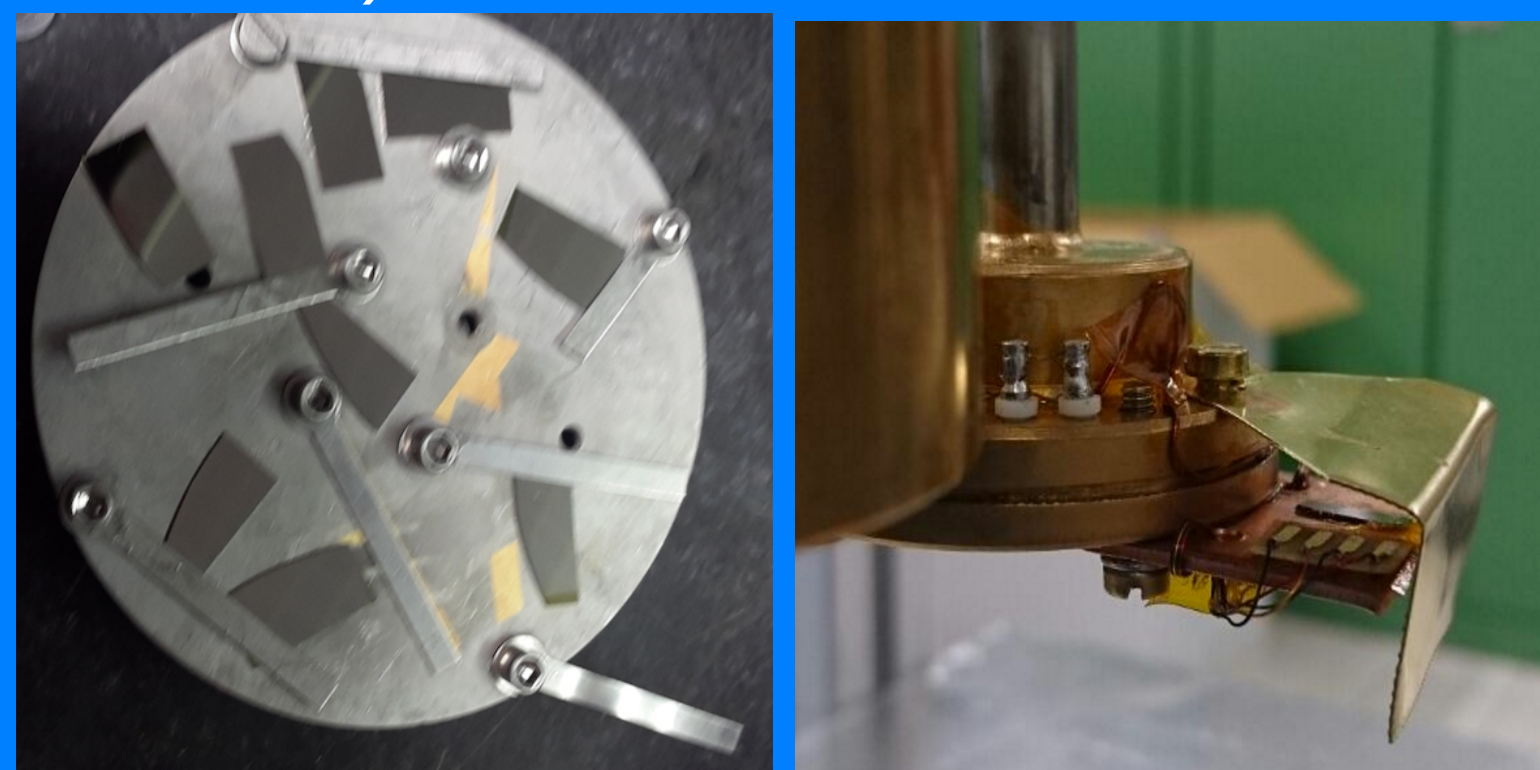


Microscope view of a spiral inductor.

Titanium T_c tuning

Titanium thin film deposition

- SiN – SiO₂ – Si clean substrate.
- substrate temperature during process less than 80 °C.
- 65 nm of Titanium deposited in 90 s using an electron gun (INFN Genova facilities).

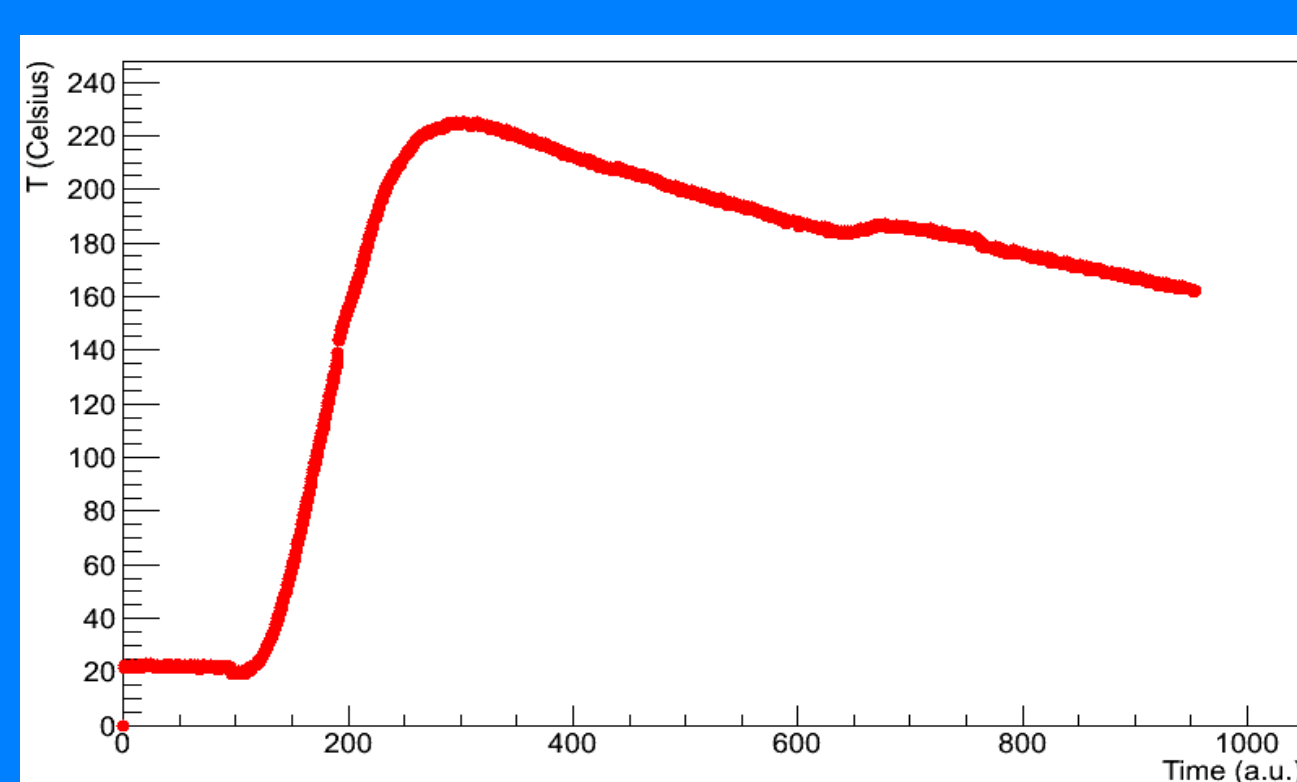


Upper left: samples after Ti deposition.

Upper right: test plate attached to the 3He cold head.

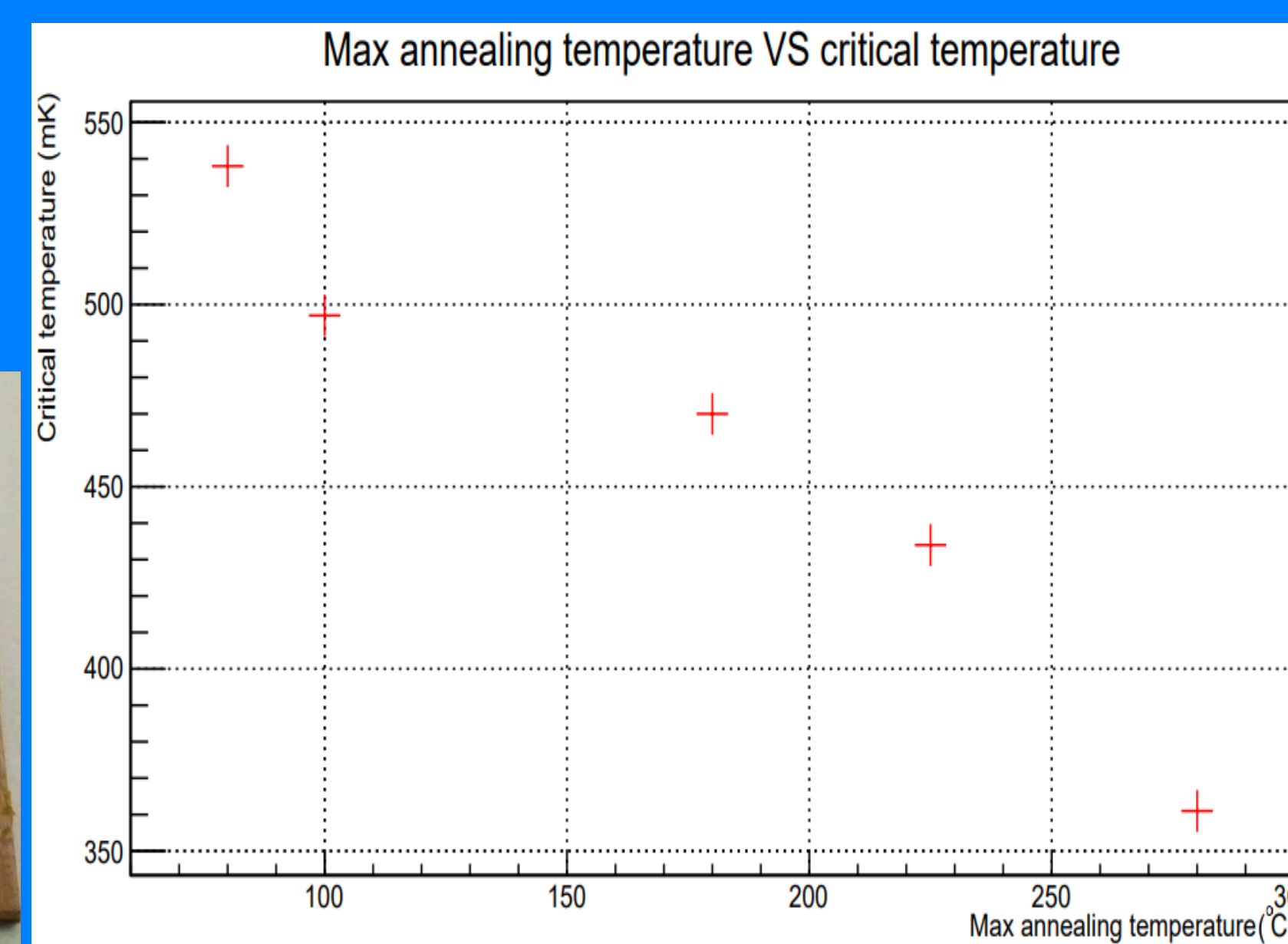
Lower left: time trend of the oven temperature during annealing.

Lower right: one of the Cu plates used for testing. The Ti strips were glued with GE7430 varnish.



Post annealing treatment

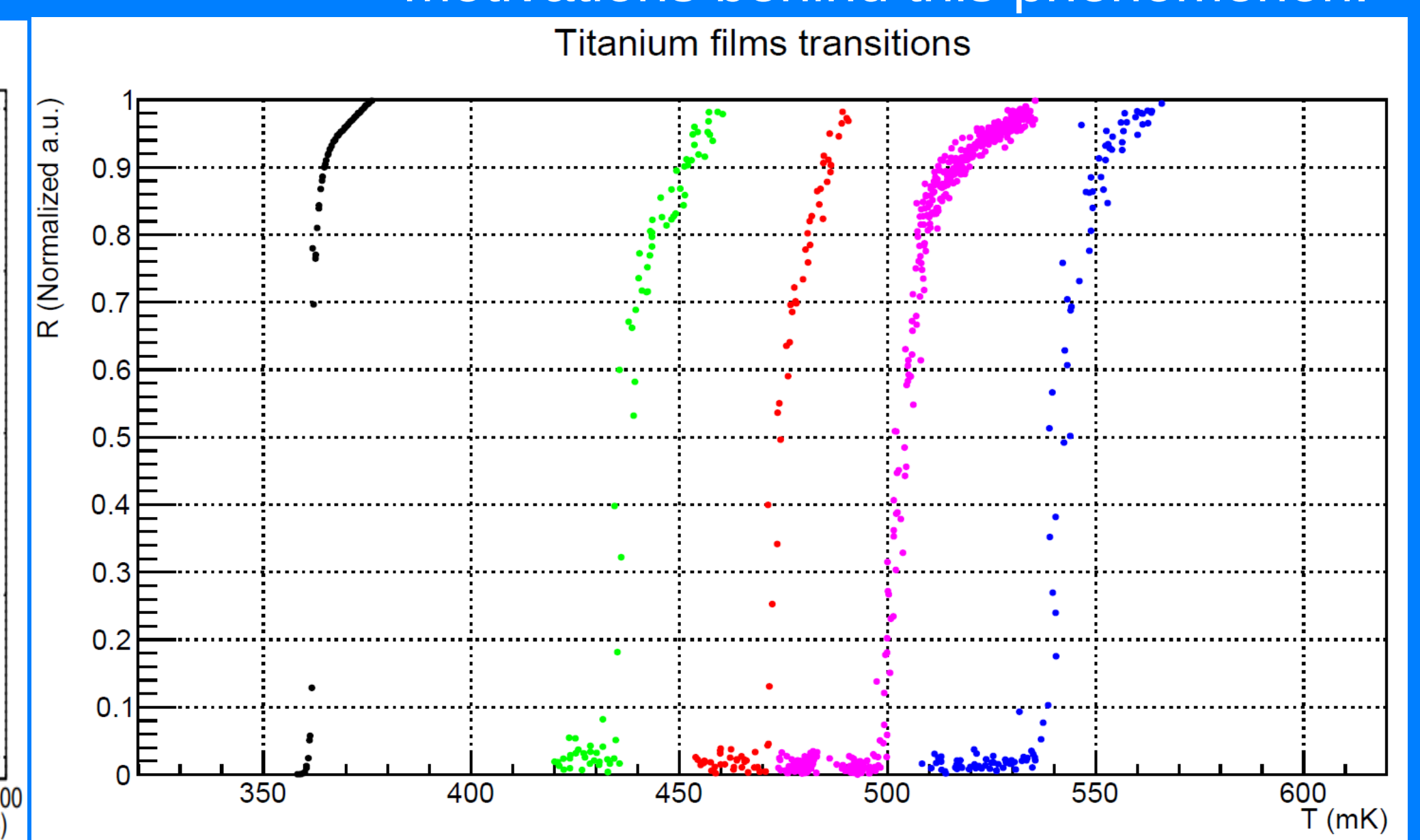
- annealing performed in an oven with Argon atmosphere to avoid Titanium oxidation (INFN Genova facilities).
- each sample is thermally treated with a different maximum temperature in the range 80 °C ÷ 270 °C.
- the entire annealing process requires 6-8 hours, including warming and slow cooling.
- Samples are cut to strips and measured in a 3He-4He sorption cooler down to ≈ 300 mK (INFN Pisa facilities). Results are reported below.



Max annealing temperature VS critical temperature

Conclusions

- Titanium T_c decreases with the increase of annealing temperature.
- a dedicated thermal process allows to set a particular T_c .
- it is possible produce TES bolometers with a T_c suitable for LSPE.
- further investigations are in progress to understand the crystallographic motivations behind this phenomenon.



Titanium films transitions

CONTACTS

davide.vaccaro@pi.infn.it; beatrice.siri@ge.infn.it

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